

Status of S-NPP VIIRS Solar and Lunar Calibration

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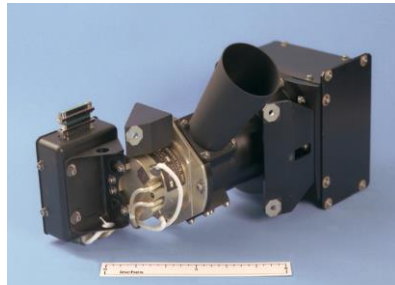
Outline

- **Solar and Lunar Calibration**
 - Strategies and Activities
 - Methodologies
- **Performance Updates**
 - On-orbit Changes and Performance Updates (Improvements)
 - Comparison of Solar and Lunar Calibration
- **Future Efforts**
- **Summary**

Solar and Lunar Calibration Strategies and Activities

15 RSB: M1-M11, I1-I3, DNB
H/L gains: M1-5 and M7
 λ : 0.4-2.3 μm

Solar Diffuser
Stability Monitor



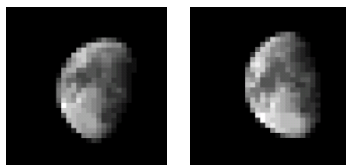
SD with a fixed screen

SD calibration each orbit

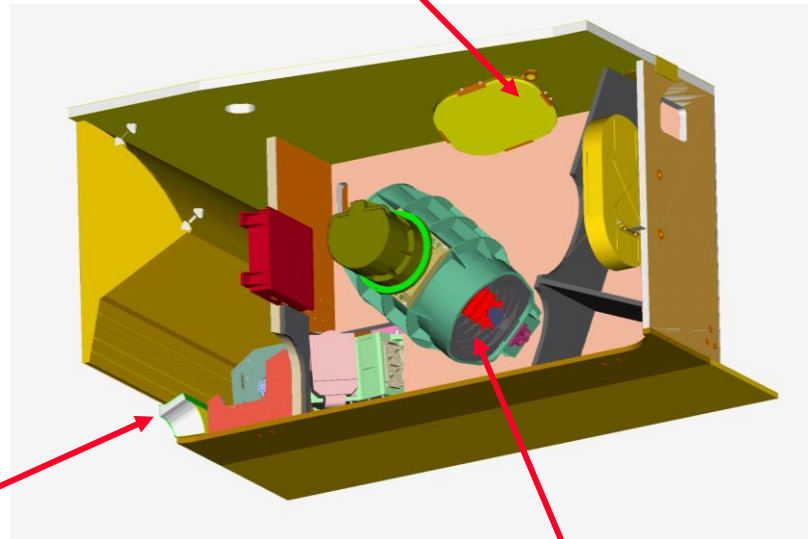
Daily operation => 3 per week
(8 min => 5 min)

Future reduction of frequency
and operation time

8-9 / year



Extended SV Port



Rotating Telescope Assembly (RTA)

Solar Calibration Methodologies

Quadratic Approach

VIIRS Radiance (L) Retrieval: $L = F \cdot L_{PL} = F \cdot (c_0 + c_1 \cdot dn + c_2 \cdot dn^2) / RVS$

F : Calibration scaling factor derived from on-orbit calibration

c_i : Pre-launch calibration coefficients (quadratic algorithm)

RVS : Sensor response versus scan-angle

Reflectance Based

VIIRS Solar Calibration: $F_{SD} = \frac{L_{SUN}}{L_{SD,PL}}$ $L_{SUN} \propto E_{SUN} \cdot BRDF(t) \cdot \tau_{SDS} \cdot \cos(\theta_{inc})$

L_{SUN} : Expected solar radiance reflected from SD panel

$L_{SD,PL}$: Retrieved solar radiance using pre-launch calibration coefficients

SD Degradation (H): $BRDF(t) = H_{Norm}(t) \cdot BRDF(t_0)$

Lunar Calibration Methodologies

VIIRS Lunar Calibration:

$$F_{MOON} = \frac{I_{ROLO}}{I_{MOON,PL}} = \frac{I_{ROLO}}{\sum_{det,sam,scan} L_{MOON,PL} \cdot \Omega_B \cdot g / N_{SCAN}}$$

I_{ROLO} : Lunar irradiance (integrated) provided by ROLO model

$I_{MOON,PL}$: Lunar irradiance retrieved using pre-launch calibration coefficients

N_{SCAN} , Ω_B , g : number of scans, pixel solid angle, aggregation factor

Ongoing and Future Activities for Lunar Model Improvements:

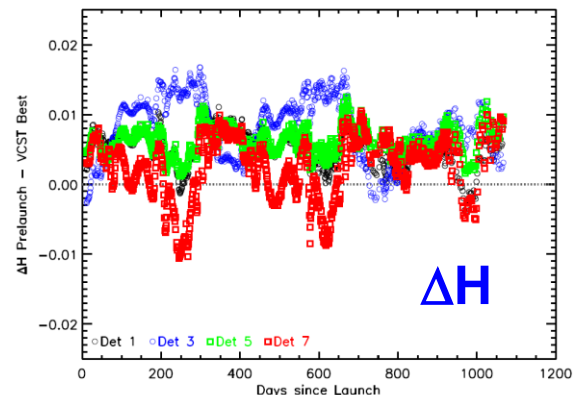
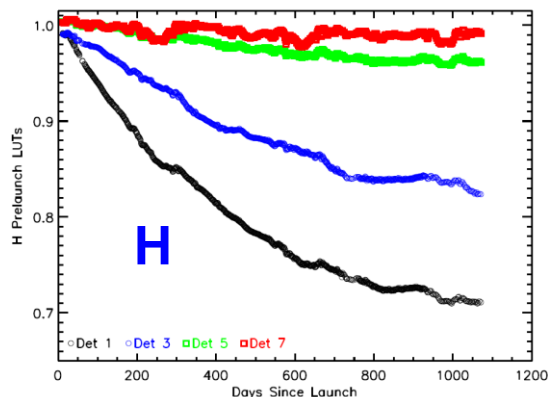
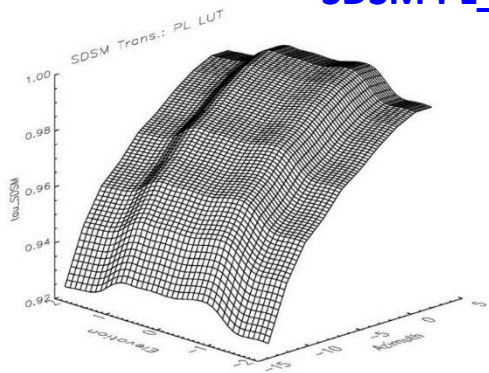
- USGS ROLO (Stone/Kieffer)
- NIST high accuracy measurements (Brown et al)
- CNES POLO data

On-orbit Changes and Updates (Improvements)

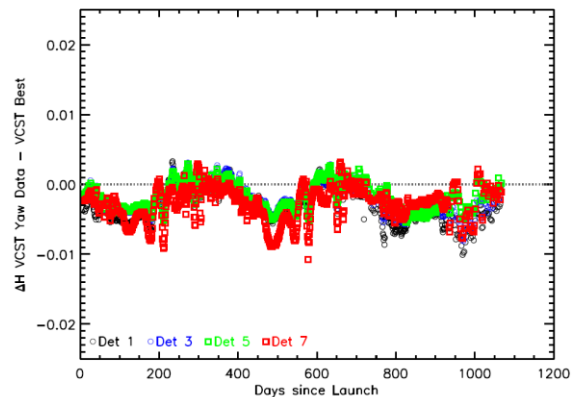
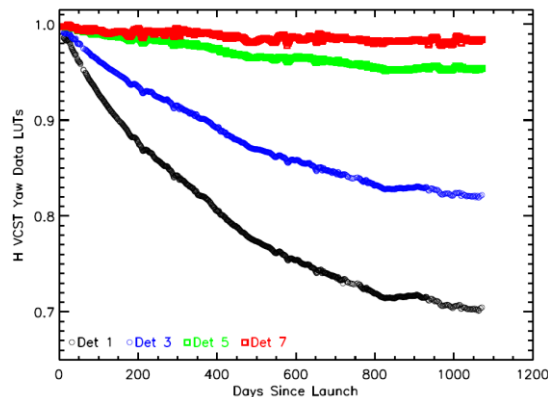
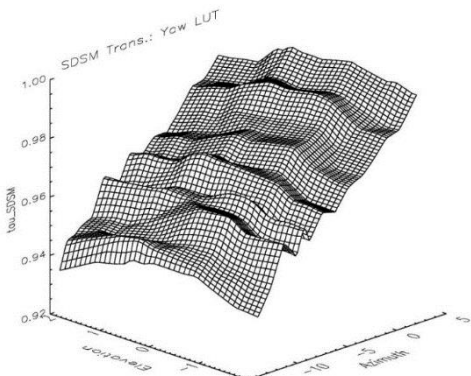
- **SD and SDSM Screen Transmission**
 - Pre-launch characterization
 - On-orbit yaw maneuvers
 - Yaw + regular on-orbit data
- **Correction for Solar Vector Error**
 - Consistently reprocessed SDR for NASA science research community
 - Different impact for VIS/NIR and SWIR
- **Modulated RSR (relative spectral response)**
 - Due to strong wavelength-dependent optics degradation
 - Different impact for solar and lunar calibration, and EV data
 - Large effect for DNB (broad bandwidth: 500-900 nm) calibration

SD and SDSM Screen Transmission (LUTs)

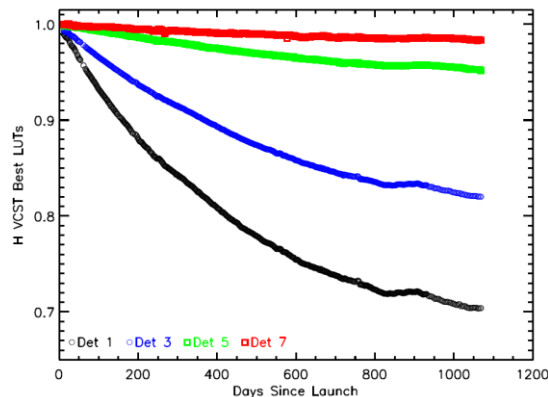
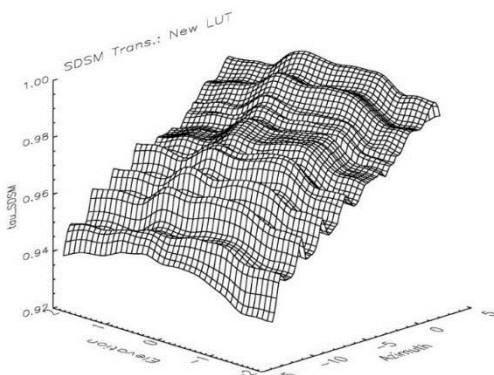
SDSM PL_LUT



SDSM Yaw_LUT



SDSM New_LUT



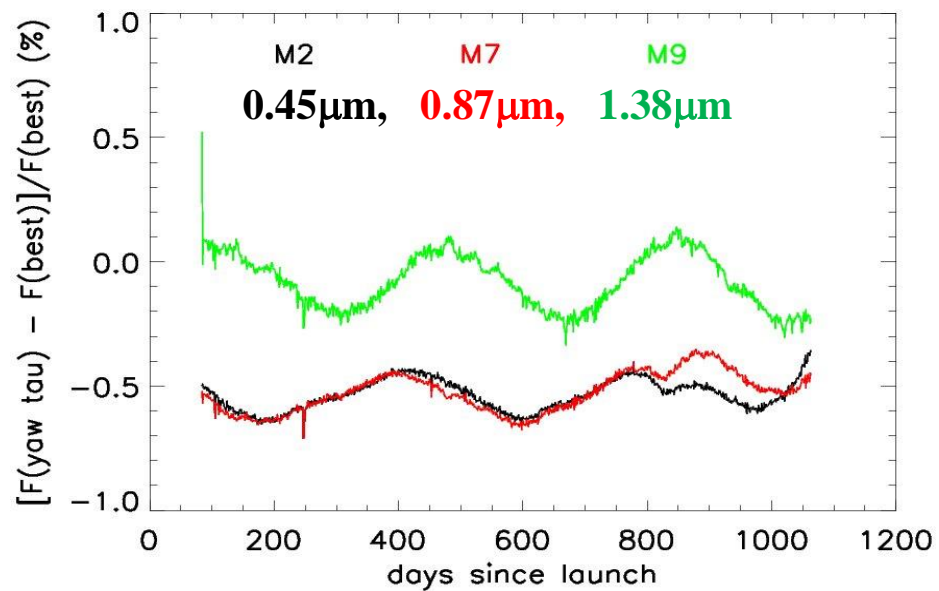
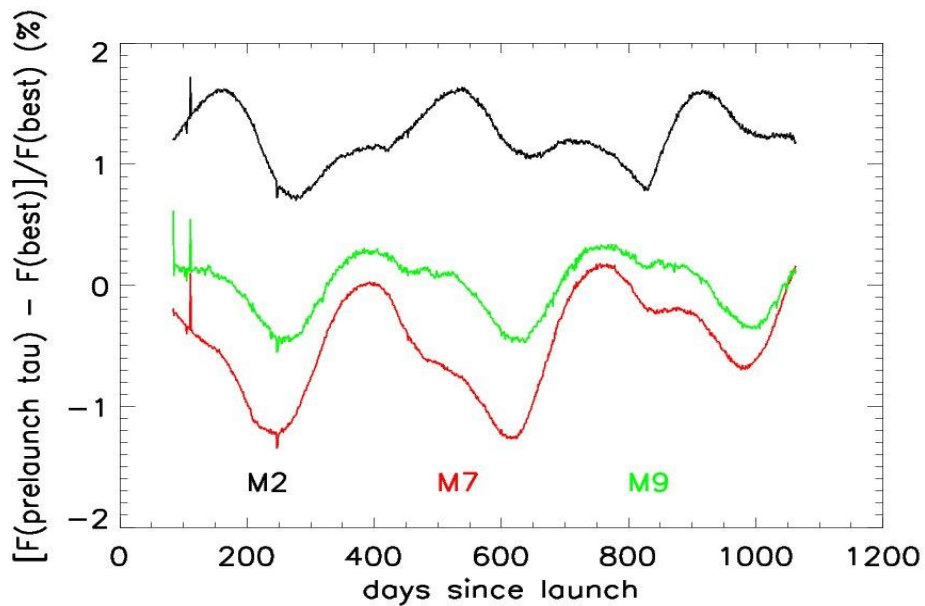
Improved H => better quality of F

Similar improvements to SD screen transmission

0.41, 0.48, 0.67, 0.86 μm

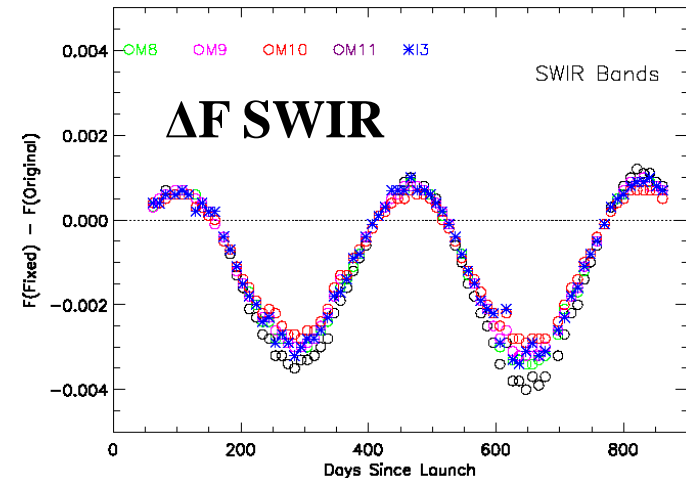
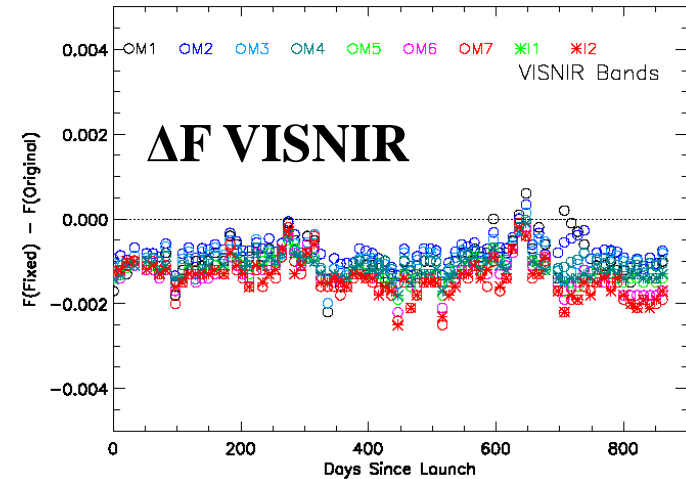
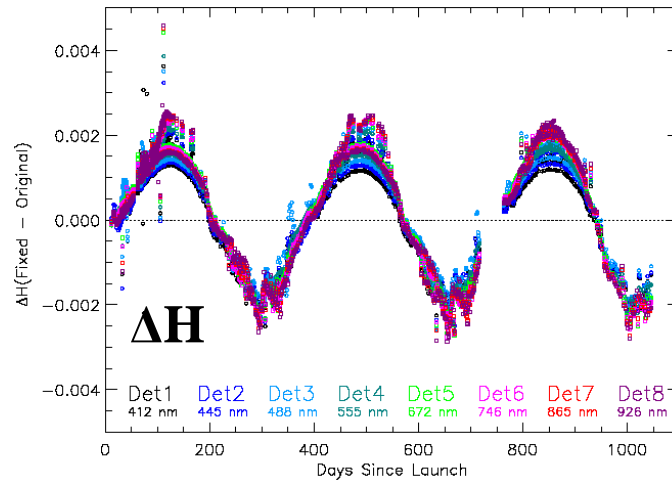
SD and SDSM Screen Transmission (LUTs)

Impact on F-factor (1/Gain)



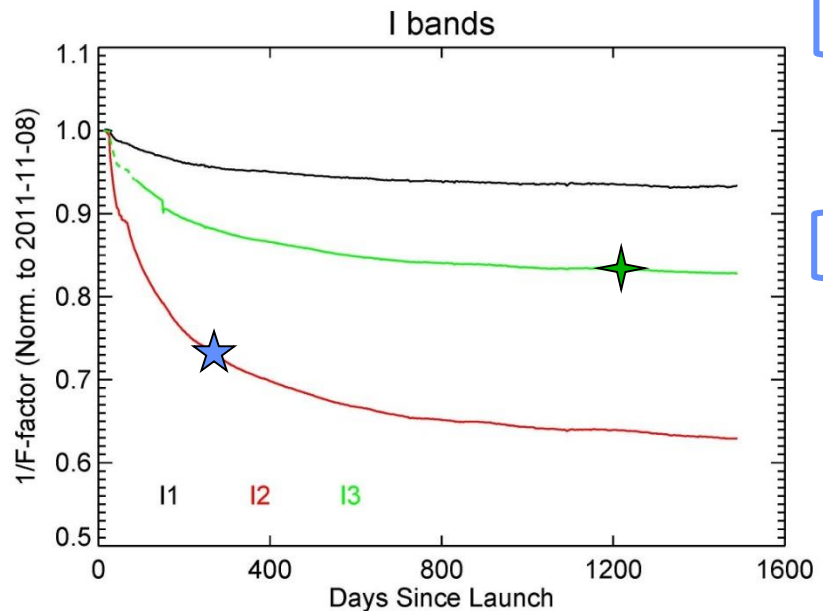
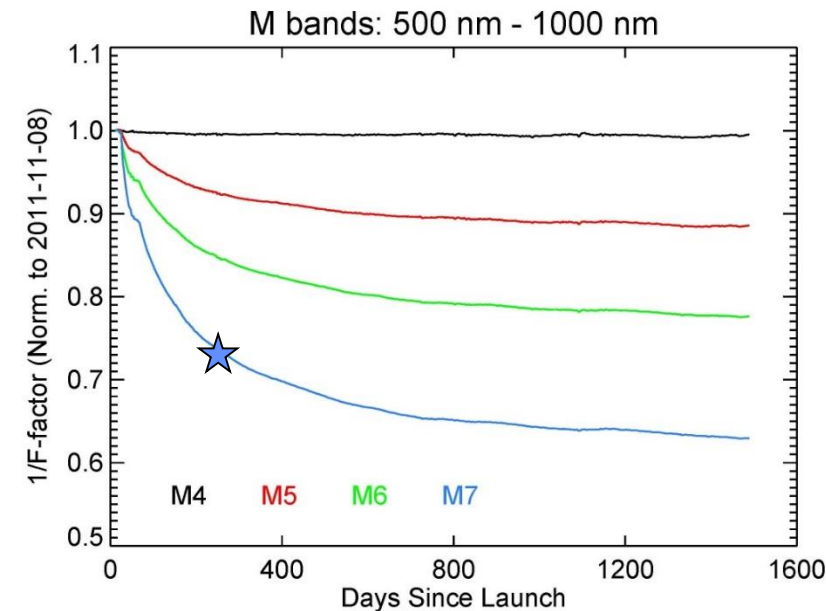
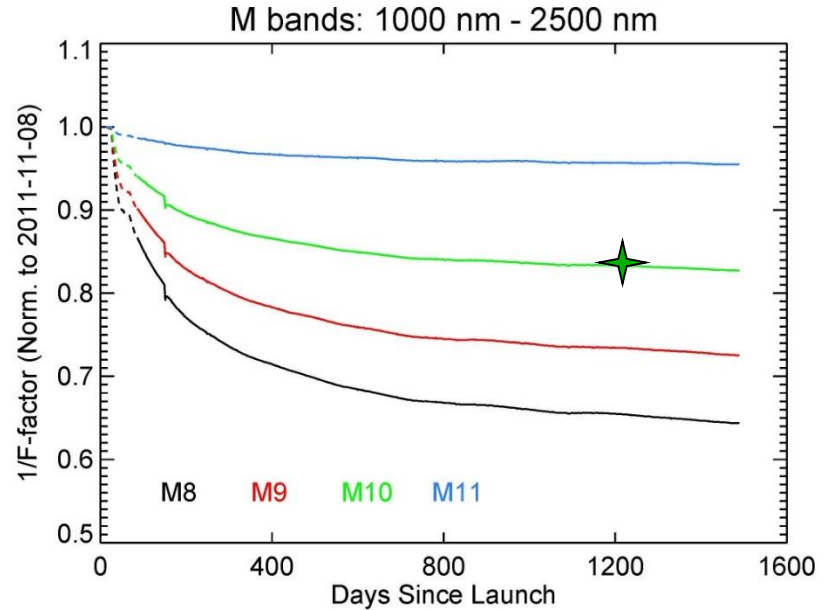
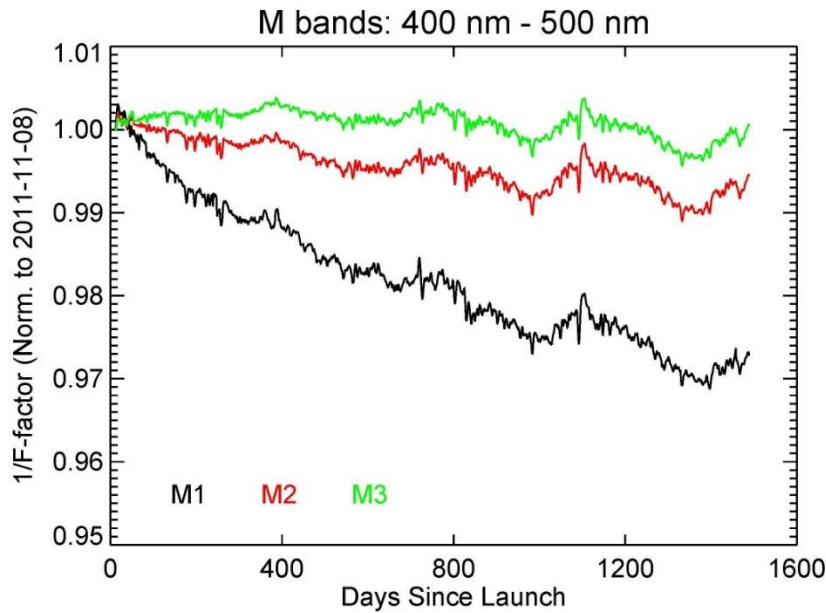
Correction for Solar Vector Error in SDR Geo Library

- A mismatch of ECI (Earth-Centered Inertial) frames when computing the transformation to spacecraft frame library leads to $\sim 0.2^\circ$ error in the solar angles used in the RSB radiometric calibration.
- Different impact for VIS/NIR and SWIR bands



- The $\cos \theta_{SD}$ factor is used in both H- and F-factor calculations.

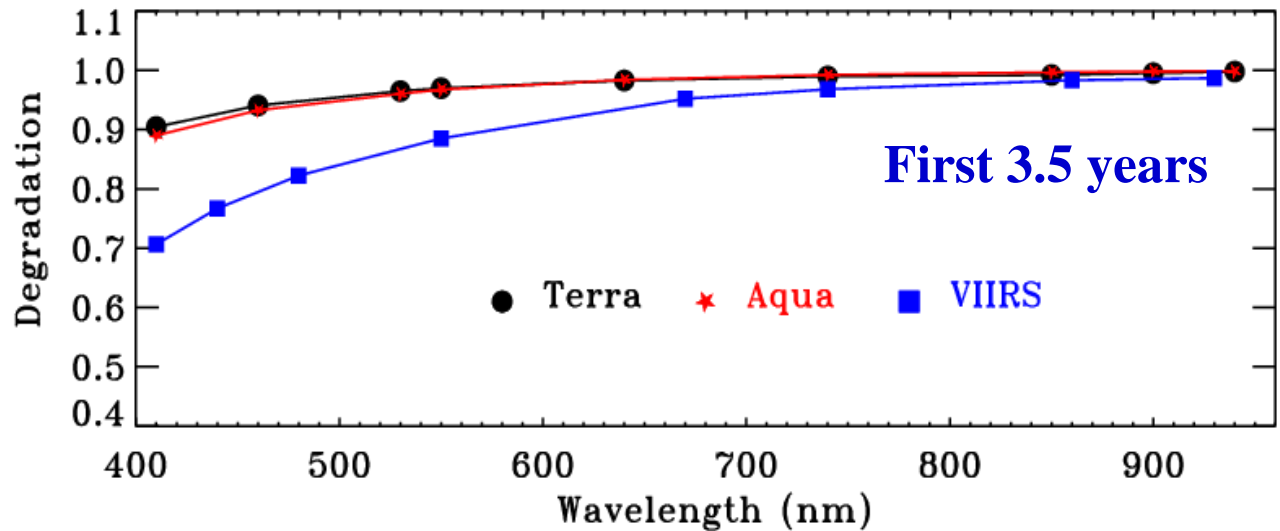
SD F-factors for VIIRS Reflective Solar Bands



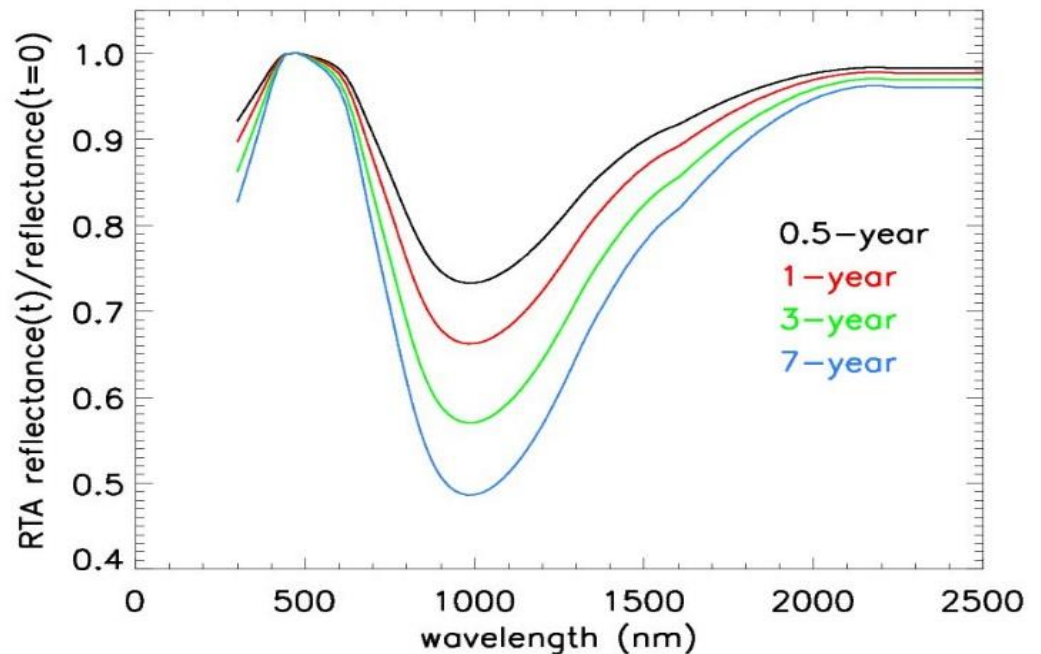
| | |
|-----|------|
| M1 | 0.41 |
| M2 | 0.45 |
| M3 | 0.49 |
| M4 | 0.56 |
| I1 | 0.64 |
| M5 | 0.67 |
| M6 | 0.75 |
| I2 | 0.87 |
| M7 | 0.87 |
| M8 | 1.24 |
| M9 | 1.38 |
| I3 | 1.61 |
| M10 | 1.61 |
| M11 | 2.25 |

Solar Diffuser and Optics Degradation

SD degradation:
Large at short
wavelength

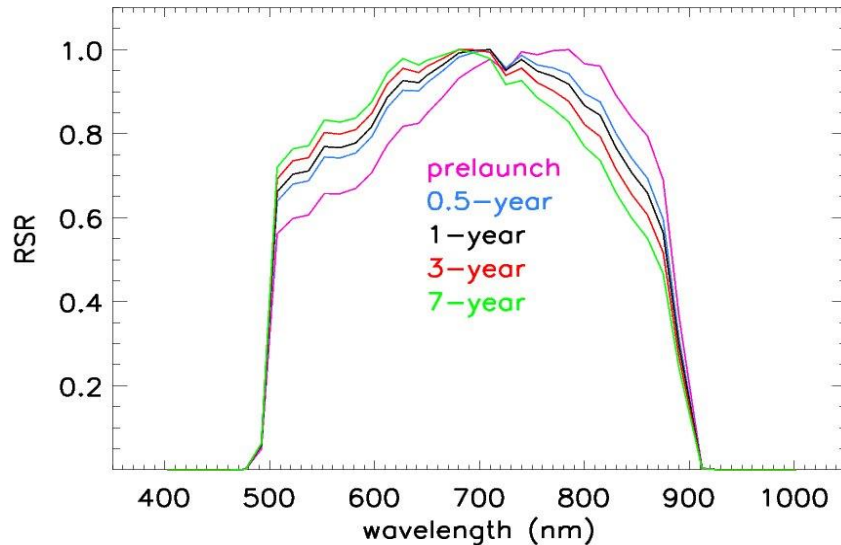


Optics degradation:
Large at NIR/SWIR



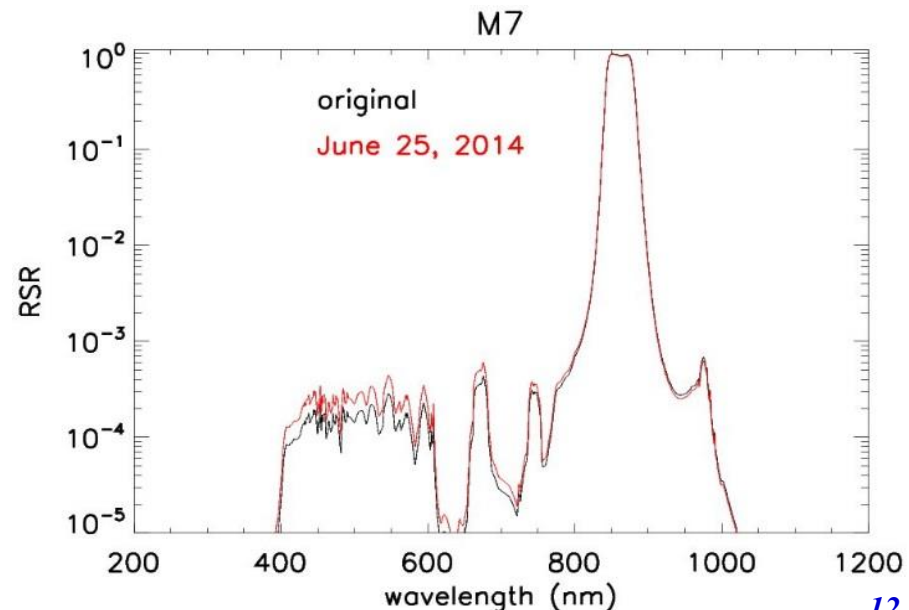
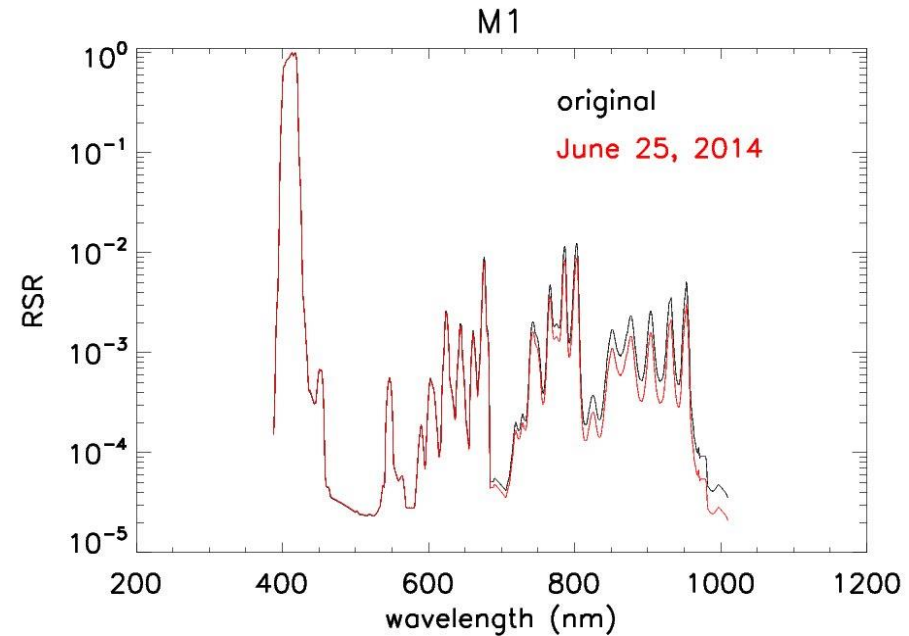
Development and Update of On-orbit Modulated RSR

Time-dependent RSR

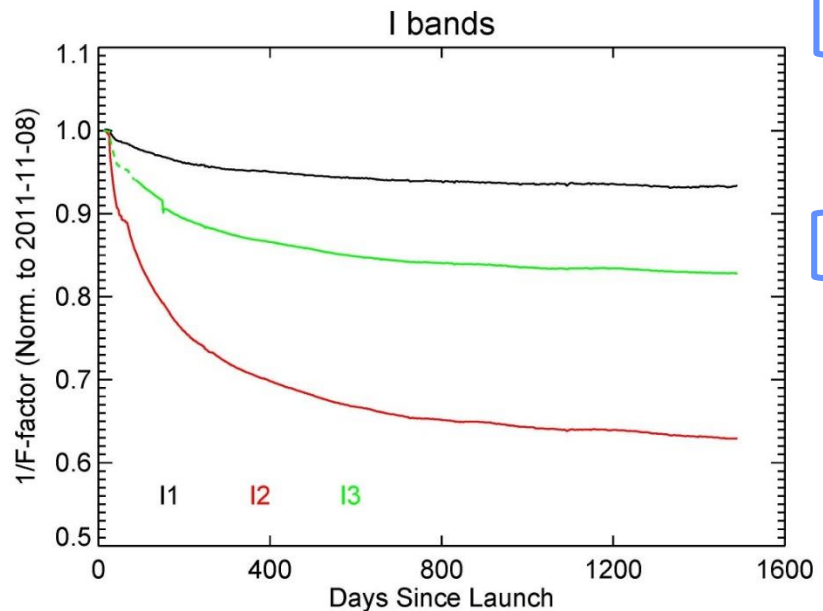
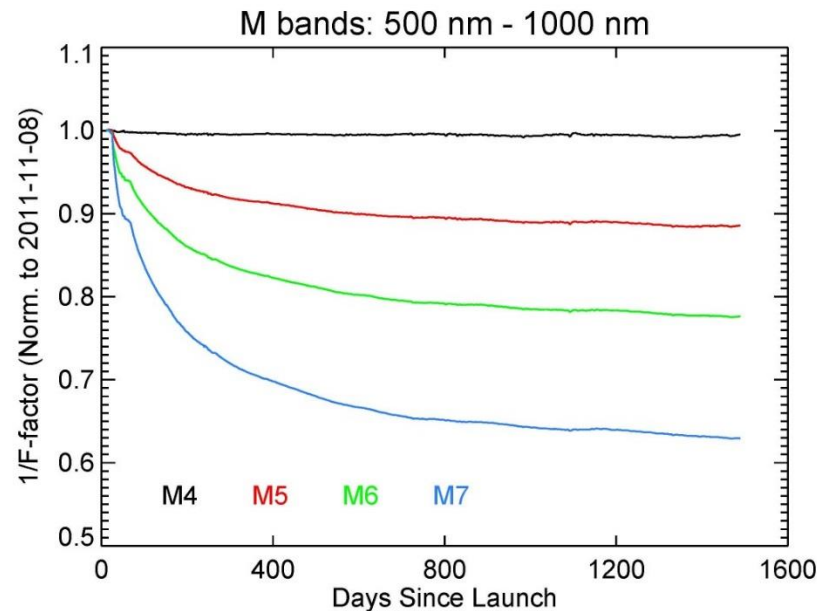
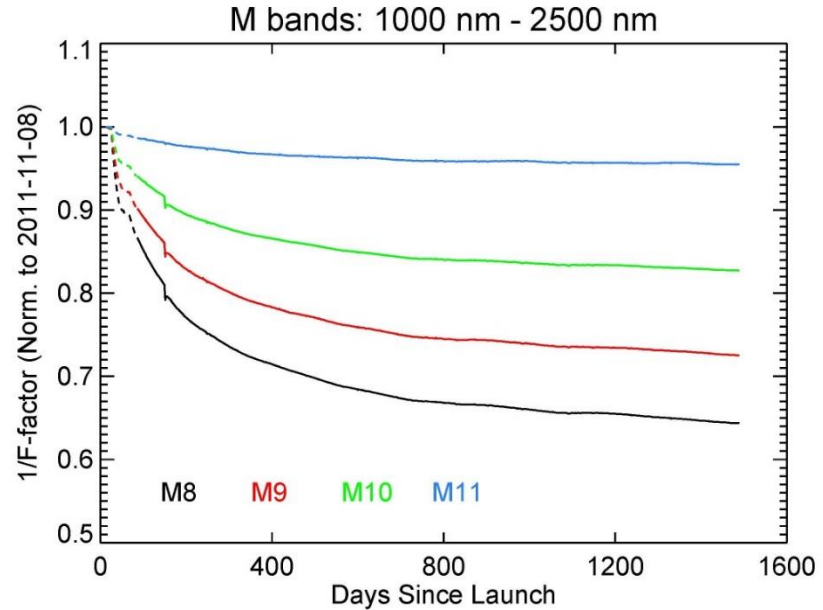
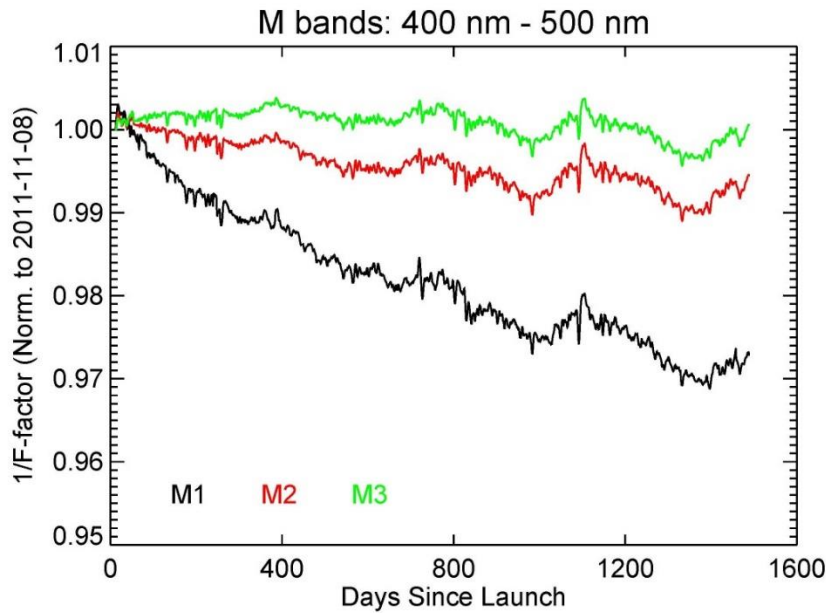


Large impact on DNB with a broad bandwidth

Small impact on bands with narrow bandwidths and non-negligible OOB responses

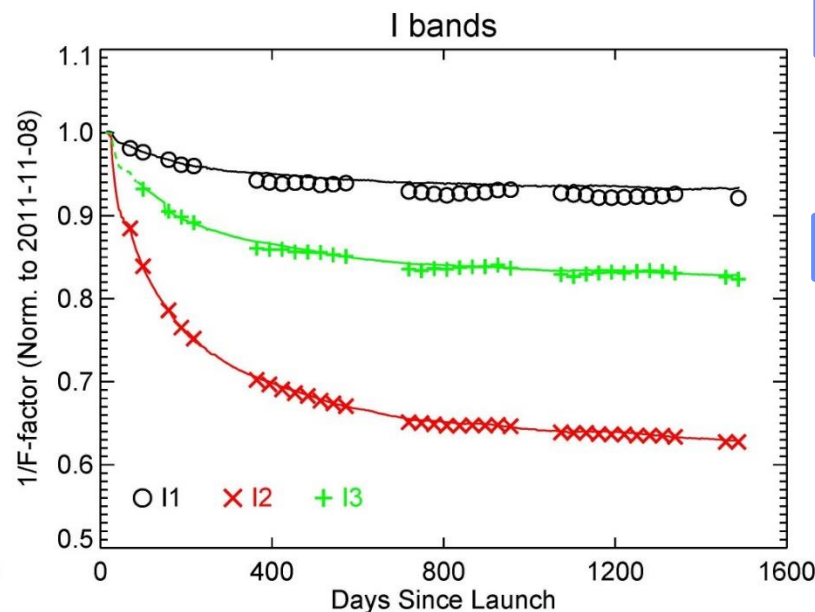
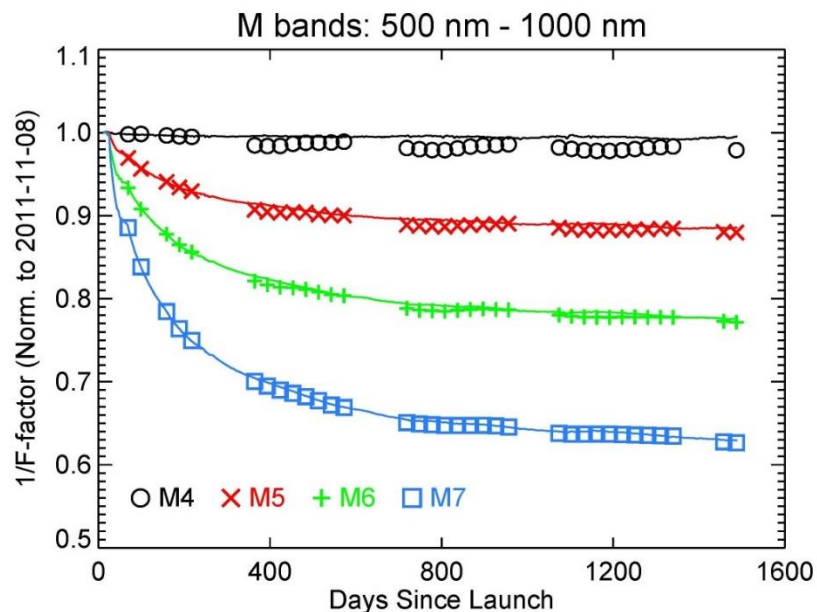
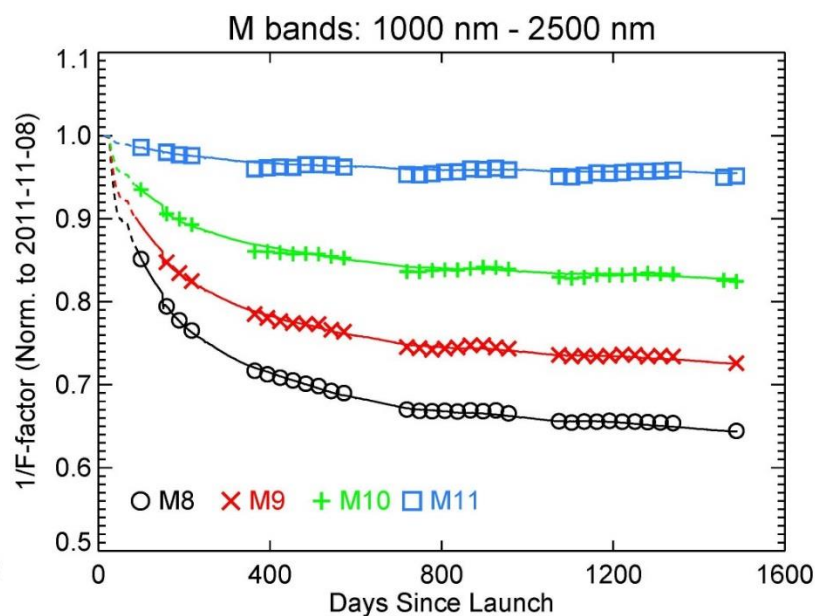
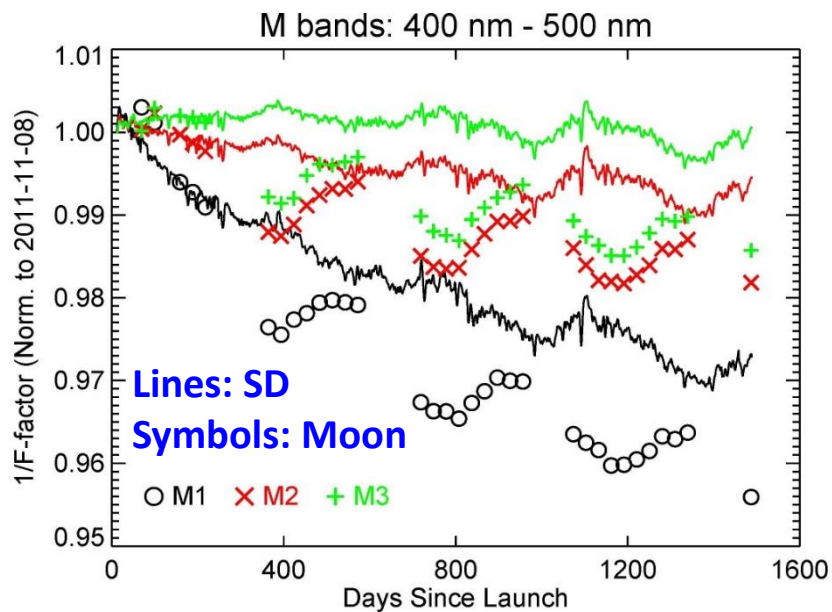


SD F-factors for VIIRS Reflective Solar Bands



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SD and Lunar F-factors for VIIRS Reflective Solar Bands

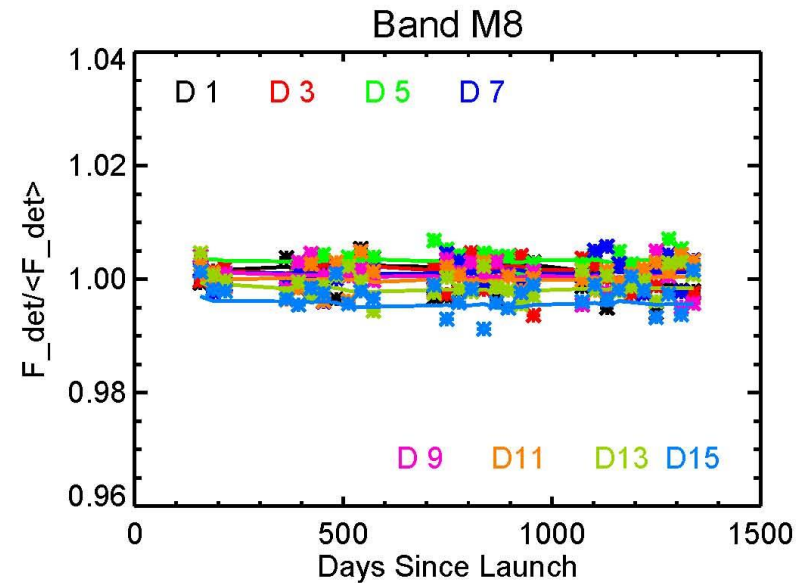
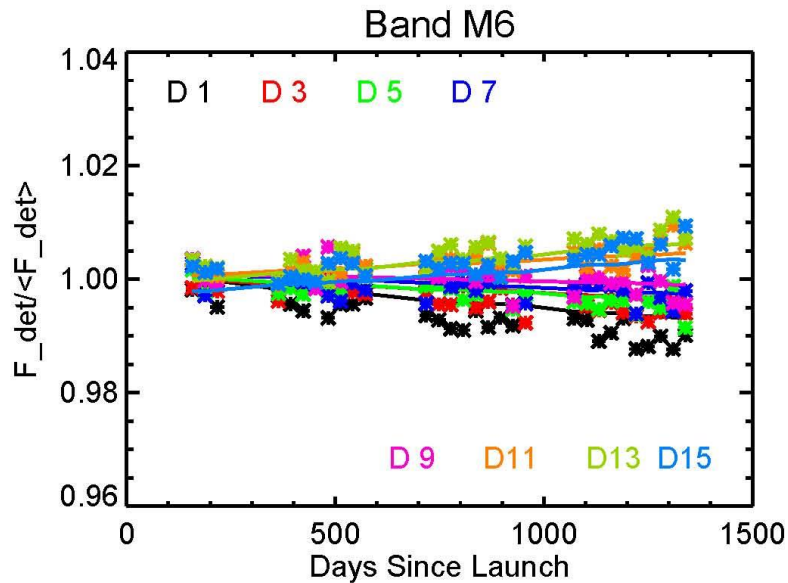
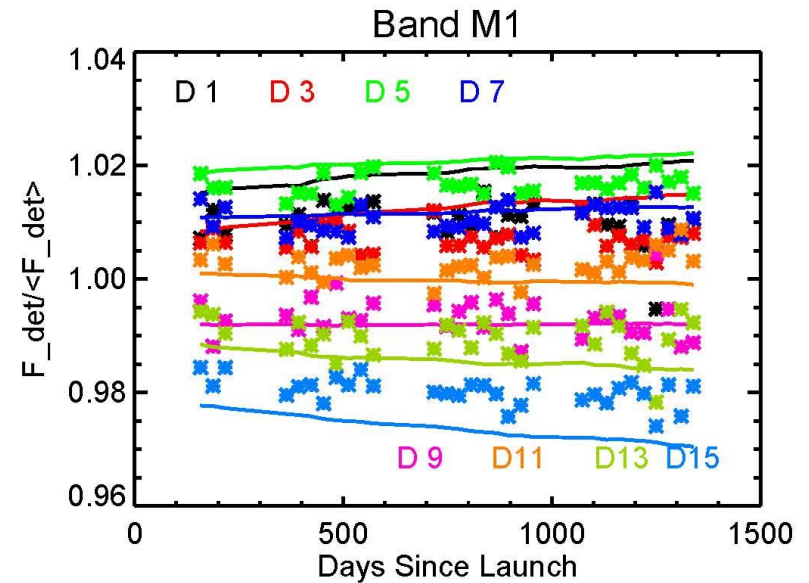
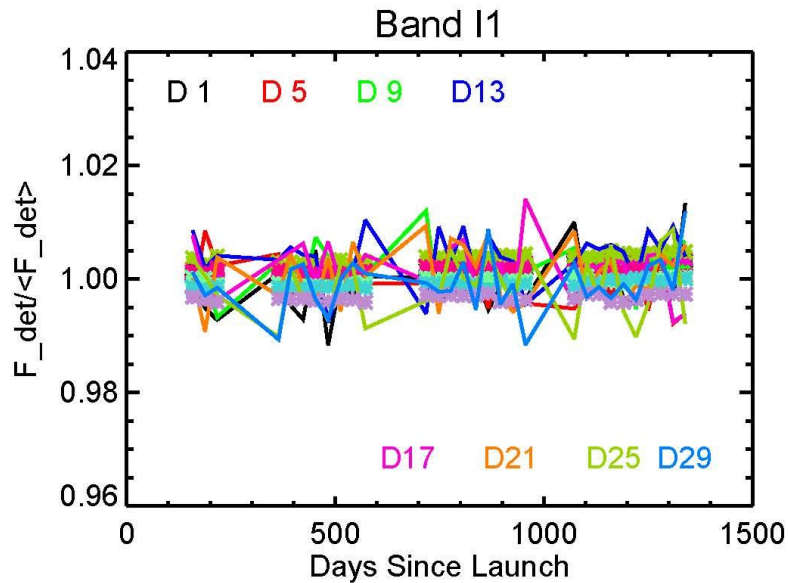


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Future Efforts

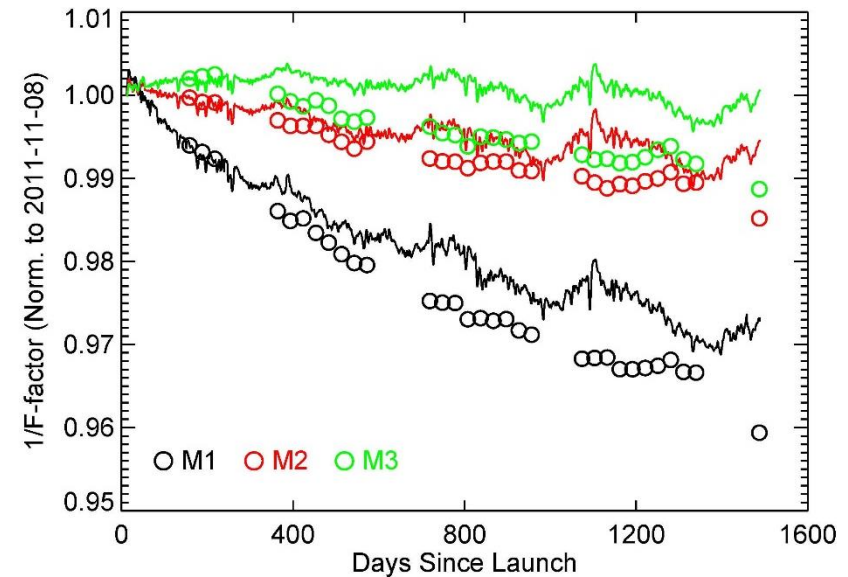
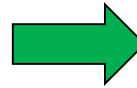
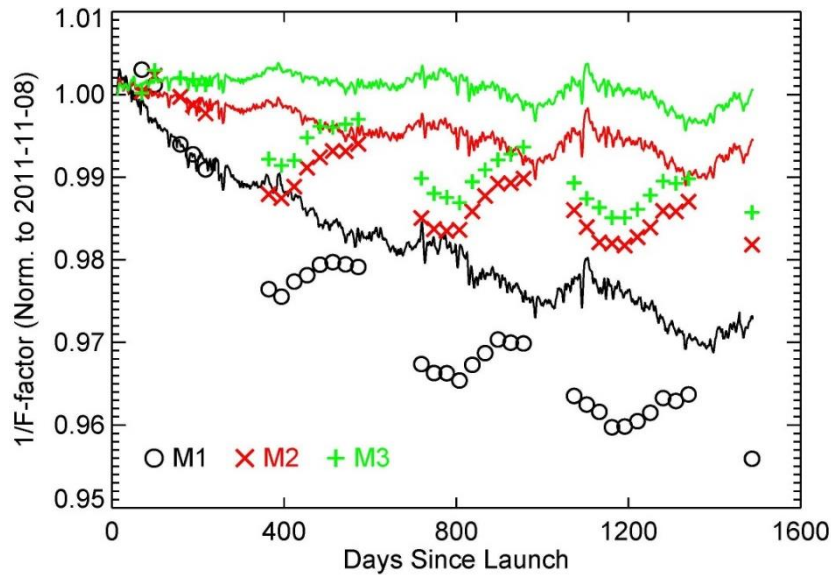
- **Combine SD and Lunar Calibration for Improved SDR LUTs**
 - SD and lunar observations are made at the same AOI
 - Remove potential impact due to SD degradation (SDSM and SD degradation uniformity)
- **Use Lunar Observations to and Characterize and Reduce Detector to Detector Calibration Differences**
 - Similar strategy developed and applied for MODIS calibration
 - Small differences in a few VIIRS spectral bands
- **Improve Lunar Calibration**
 - Absolute - effort by NIST/USGS (goal: 0.5%) and by GSICS/USGS (goal: 2%)
 - Relative - response trending and calibration inter-comparison

Detector to Detector Calibration Differences

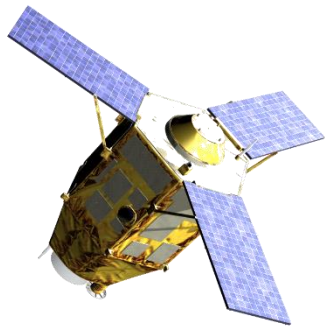


Approaches for Lunar Calibration Improvements

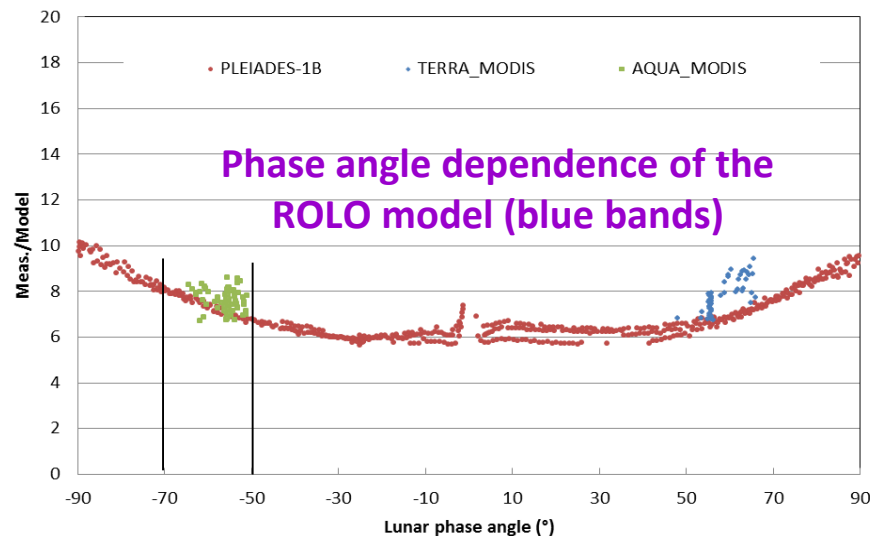
With an empirical libration correction



Impact due to lunar phase angles



Pleiades: POLO



Phase angle dependence of the
ROLO model (blue bands)



Summary

- **S-NPP VIIRS continues to perform well, meeting the need for operational users (SDRs/EDRs from IDPS) and science community (reprocessed SDRs/EDRs)**
 - NASA VCST and SIPS effort
 - NOAA reprocessing plan
- **Improved understanding of both solar and lunar calibration led to generation of consistent LUTs and high-quality data products**
- **Future efforts planned to address various challenging issues**
 - Near-term
 - Long-term